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11.1. Hom-1Imanr Maves fa a Tro-cempenent Mot Plasma.

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1. 

## IUTRODITTION

Following clemmow ${ }^{(1,2)}$, ve studv nonlinear wave propagation in a hot, collisionless nlasma consistina of niectr and lons. No assume that the nlasma is un!onnded and that there is no ambient magnetic figid. The modol user is the BnltamannVlasov equation ( $B-V$ ecuations) in a Lnrentz frame of referenc $S$ in which tho snace-denendcnce in rliminated ${ }^{(3)}$.

Te invectiaate transverfe wavna for two cascs, irine norturbation technicue: (i) When the niama in not extromoly hot so that the temnerature effect can be treator so a small correction to the cold mlama caso and the amnlitude of the wave is larae. (ii) Ther the ware amplitudr is small and that there is no reetriction on the temnerature. In hoth cases, tho dispersion relationsare shtainer and the resulte are nresenter in ? way that the dectun and ion fffecte stand out scmarately

The plan of tho namer ies Gectinn 2 nresonts a auncr formulation of the prohlem. Soctinn ? snecialises to transurs propagation and dovelon the marter equations (15) and (le). In sub-section 3.1 wicord the results for the onld plosma. The main rosults of this narer are in suh-sections 3.2 and 3.3 describing the ajenersion rnlations for strone maves (i.e. Inr amniture) with first orfre temnorature fffect and for weak w(i.c. small amilituo) with first order non-linear correction leeping temnerature arbitrarv.

We conaider $S^{\prime}$ as the laboratory frame in which the velocity of the wave is $(n, n, c / n)$ and $S$ the frame in which there is no snece-dependence and which is moving with velocity (0, n, ne) relative to $s^{\prime}$ ( $n$ heine the refractive index of the medium). Alj nur calculations will he in frame s which can thon be transformen to frame $G^{\prime}$ with the helr of a Lorenta transformation.

Dun to the shscnce of the snatial denendance of the
ficl.ds in frame $s$, Maxicll ecuations imnly that the mannetic ficld $B$ is constant ane that the number densitior of electrons and ions are ecual, sav N. Furtionr, the curl ne $B$ orvation ir roduced to

$$
\begin{equation*}
-\varepsilon_{0} \dot{\dot{E}}=\sum_{\alpha=e, j}{ }^{I}-\alpha \tag{1}
\end{equation*}
$$

We consicor the snecial case $Z=0$. Then the relativistic B-V eduations for electrons and lons will be

$$
\begin{equation*}
\frac{\partial f_{\alpha}}{\partial t}+\frac{ๆ_{\alpha}}{m_{\alpha} c} E \cdot \frac{\partial f_{\alpha}}{\partial \underline{u}_{\alpha}}=0 \tag{2.}
\end{equation*}
$$

Where $\underline{u}_{\epsilon}$ and $\underline{u}_{\mathrm{i}}$ are the recuced velocities of electrone and ions respectively, defined in terms of the ordinary velocition $\underline{v}_{e}$ and $\underline{v}_{i}$ by

$$
\underline{u}_{\alpha}=\frac{r_{\alpha} \underline{v}_{\alpha}}{c}, \quad r_{\alpha}=\left(1+\underline{u}_{\alpha}^{2}\right)^{\frac{1}{2}}
$$

Also $N f_{\alpha}\left(\underline{u}_{v}, t\right)$ is the distribution function. Now. using $\underline{E}=-\underline{\underline{E}}$ and defining

$$
\lambda_{\alpha}=-\frac{q_{\alpha}}{m_{\alpha} c} \quad \text { where } q_{i}=t e \text { and } q_{e}=-e
$$

the $B-V$ equation may be expressed as

$$
\begin{equation*}
\frac{\partial f_{\alpha}}{\partial t}+\lambda_{\alpha} \dot{\mathbb{A}} \cdot \frac{\partial_{\alpha}}{\partial u_{\alpha}}=0 \tag{3}
\end{equation*}
$$

These equations have a general solution

$$
\begin{equation*}
f_{\alpha}\left(\underline{u}_{\alpha}, t\right)=F_{\alpha}\left(\underline{u}_{\alpha}-\lambda_{\alpha} \underline{A}\right) \tag{1}
\end{equation*}
$$

where $F_{\alpha}$ is an arbitrary function of its argument; ${\underset{\sim}{-\alpha}}-\lambda_{\alpha}$ is $\frac{1}{m_{\alpha} c}$ times the renerglizec momentum. Now

$$
\begin{aligned}
J_{\alpha} & =N r_{\alpha} c \int_{\alpha}^{\frac{u}{\alpha}} e_{\alpha i}\left(\underline{u}_{-i}, t\right) d^{3} \underline{u}_{\sim i} \\
& =\frac{N c_{\alpha}^{c}}{\lambda_{\alpha}} \frac{\partial V}{\partial \underline{\lambda}}
\end{aligned}
$$

where

$$
\begin{equation*}
v_{\alpha}=\int\left[1+\left(\underline{11}_{\alpha}+\lambda_{\alpha} \underline{\lambda}\right)^{2}\right]^{\frac{1}{2}} F_{\alpha}\left(\underline{u}_{\alpha}\right) d^{3} \underline{u}_{\alpha} . \tag{5}
\end{equation*}
$$

The equation (1) may therefore be rewritten as

$$
\begin{equation*}
\ddot{\ddot{A}}+\sum_{\alpha} \frac{\omega_{\alpha}^{2}}{\lambda_{\alpha}^{2}} \frac{\partial V}{\partial \underline{A}}=0 \tag{}
\end{equation*}
$$

where

$$
\omega_{\alpha}^{2}=\frac{N q_{\alpha}^{2}}{\varepsilon_{-} m_{\alpha}}
$$

A is also a constant.. With constant $A$, the equation (12) under appropriate initial conditions has a solution

$$
\begin{equation*}
\Phi=\omega t ; \omega=\frac{h}{A^{2}} \tag{13}
\end{equation*}
$$

and the equation (10) reduces to the form

$$
\begin{equation*}
\frac{\partial V}{\partial A}=\frac{h^{2}}{A^{3}} \text { ox } \frac{1}{A} \frac{\partial V}{\partial A}=\omega^{2} \tag{14}
\end{equation*}
$$

There is thus in $S$ frame a monochromatic circularly polarized field of vector potential

$$
\underline{I}=\left[A \cos (\omega t), A \sin (\omega t), \Lambda_{z}\right]
$$

and

$$
\underline{E}=A \omega[\sin (\omega t),-\cos (\omega t), 0]
$$

where $A, A_{z}$ and $w$ satisfy the equations

$$
\begin{align*}
& \sum_{\alpha} \frac{\omega_{\alpha}^{2}}{\lambda_{\alpha}^{2}} \frac{\partial V_{\alpha}}{\partial Z_{z}}=0  \tag{15}\\
& \sum_{\alpha} \frac{\omega}{\alpha}_{\lambda_{\alpha}^{2}}^{\lambda_{\alpha}^{2}} \frac{\partial V_{\alpha}}{\partial A}=\omega^{2} \tag{1.6}
\end{align*}
$$

Transforming the results to the laboratory frame $S^{\prime}$ again yields a purely transverse circularly nolarized wave with velocity $(0,0, c / n)$ and ancuilar frequency $\omega^{\prime}$. The fields in the laboraio frame $s^{\circ}$ will be

$$
\underline{E}^{\prime}=E_{\cap}^{\prime}\left\{\sin \left[\omega^{\prime}\left(\theta^{\prime}-n z^{\prime} / c\right)\right],-\cos \left[\omega^{\prime}\left(t^{\prime}-n z^{\prime} / c\right)\right], 0\right.
$$

and

$$
\underline{B}^{\prime}=\frac{n}{c} \dot{\underline{n}} \times E^{\prime}
$$

where the electric field amplitudes in $S$ and $S$, are related by

$$
\begin{equation*}
\frac{E_{0}}{\omega}=\frac{E_{0}^{\prime}}{\omega^{\prime}}=A \tag{17}
\end{equation*}
$$

The dispersion relation is obtained by determining $\omega$ in term., of $A$ from (15) and (16) and then substituting it in

$$
\begin{equation*}
\omega=\left(1-n^{2}\right)^{\frac{1}{2}} \omega^{\prime} \tag{18}
\end{equation*}
$$

3.1

Disnersion Relation in Cold plasma

The cold nlasna results can ho obtained by taking anisotropic streaming distributions, inc.

$$
\begin{equation*}
F_{\alpha}\left(\underline{u}_{\alpha}\right)=\delta\left(F_{\alpha}\right) \delta\left(\eta_{\alpha}\right) \delta\left(\zeta_{\alpha}-u_{\alpha \cap}\right) \tag{19}
\end{equation*}
$$

where

$$
\underline{u}_{\alpha}=\left(\xi_{\alpha}, \eta_{\alpha}, \zeta_{\alpha}\right)=\left(\rho_{\alpha} \cos \phi_{\alpha}, \rho_{\alpha} \cos \phi_{\alpha}, \zeta_{\alpha}\right)
$$

and $u_{\alpha o}$ is the reduced streaming velocity given by

$$
\begin{equation*}
u_{\alpha n}=\frac{v_{\alpha 0}}{c}\left(1-\frac{v_{\alpha n}^{2}}{c^{2}}\right)^{-\frac{1}{2}}=\frac{\gamma_{x 0}{ }^{\gamma} \alpha 0}{c} . \tag{21}
\end{equation*}
$$

The velocities $u_{e o}$ and $u_{i o}$ are related through the momentum conservation equation as

$$
\begin{equation*}
u_{i o}+\mu u_{e o}=u_{0} \text { (constant), } \mu=\frac{m_{e}}{m_{i}} \tag{2.2}
\end{equation*}
$$

The function $V_{\alpha}$ now tales a simpler form

$$
\begin{equation*}
v_{\alpha}=\left[1+\lambda_{\alpha}^{2} A^{2}+\left(u_{\alpha 0}+\lambda_{\alpha} A_{z}\right)^{2}\right]^{\frac{3}{2}} \equiv \Lambda_{\alpha} \text { (say) } \tag{23}
\end{equation*}
$$

so that

$$
\begin{equation*}
\frac{1}{\lambda_{\alpha}} \frac{\partial V_{\alpha}}{\partial A_{z}}=\frac{L_{\alpha}}{\Delta_{\alpha}} \quad ; \quad L_{\alpha}=n_{\alpha 0}+\lambda_{\alpha} A_{z} \tag{24}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{1}{\lambda_{\alpha_{1}}^{2} A} \frac{\partial V_{\alpha}}{\partial A}=\frac{1}{\Delta_{\alpha}} \tag{25}
\end{equation*}
$$

How observing $\omega_{i}^{2}=\mu \omega_{e}^{2}$ and $\lambda_{i}=-\mu \lambda_{e}$, and using mentations (21) ant (25), the equations (15) and (15) become

$$
\begin{equation*}
\frac{L_{e}}{\Lambda_{n}}-\frac{L_{i}}{\Delta_{i}}=0 \tag{f}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{1}{\Delta_{e}}+\frac{\mu}{\Delta_{i}}=\frac{\omega^{2}}{\omega_{e}^{2}} \tag{.27}
\end{equation*}
$$

Also

$$
\begin{equation*}
L_{\alpha}=\left(1+i_{\alpha}^{2} A^{2}\right)^{\frac{1}{2}} \Omega \tag{28}
\end{equation*}
$$

where

$$
\begin{equation*}
\Omega=\frac{u_{0}}{\left(1+\lambda_{1}^{3} A^{2}\right)^{\frac{1}{2}}+\mu\left(1+\lambda_{0}^{2} A^{2}\right)^{\frac{1}{2}}} \tag{29}
\end{equation*}
$$

Therefore

$$
\begin{align*}
\Delta_{\alpha} & =\left(1+\lambda_{\alpha}^{2} \Lambda^{2}+L_{\alpha}^{2}\right)^{\frac{1}{3}} \\
& =\left(1+\lambda_{\alpha}^{2} \Lambda^{2}\right)^{1!}\left(1+\Gamma^{2}\right)^{\frac{1}{2}} \tag{3n}
\end{align*}
$$

Thue the disnersion rolation (27) 'eecomes

$$
\begin{equation*}
\frac{\omega^{2}}{\omega_{e}^{2}}=\frac{1}{\left(1+\Omega^{2}\right)^{1 / 2}}\left(\frac{1}{\left(1+\lambda_{\Theta}^{2} A^{2}\right)}+\frac{\mu}{\left(1+\lambda_{i}^{2} \Lambda^{2}\right)^{1 / 2}}\right) \tag{31}
\end{equation*}
$$

Note that the innic contrihution which anmars as an adritive term can he sionifieant, penecially when the amnlitude of the wave is laree.
3.2

Disnorsion Relation in fot Dlasma
(First Order Tomnorature Corroctinn)
Unless the plasma ie fxtrrmely hot, we may use a nerturbation technime to calculate firat order tomnerature correctinn to the cold nlasma resilt of the provious section. To do that, we first transform the cartcoinn varinbles of intearation in the exprossion of $v_{\alpha}$ to the fxame $s_{\alpha}$ which is
moving with velncity $\left(\cap, n, v_{\alpha 0}\right)$ relative to , and then expant the interrand as a nower ecrics. Th firct order correction is obtained by truncatine tho scrics at the nuadratic torms.

Thre Jorentz traneformations ar

$$
\xi_{\alpha}^{\prime \prime}=\xi_{\alpha}, \eta_{\alpha}^{\prime \prime}=\eta_{\alpha}, \zeta_{\alpha}^{\prime \prime}=\gamma_{\alpha \cap}\left[\zeta_{\alpha}-\frac{v_{\alpha o}}{c} \gamma_{\alpha}\right]
$$

whrire

$$
\gamma_{\alpha 0}=\left(1-\frac{v_{\alpha 0}^{2}}{c^{2}}\right)^{-\cdots \frac{1}{k}}
$$

and

$$
\gamma_{\alpha}^{\prime \prime}=\gamma_{\alpha n}\left(r_{\alpha}-\frac{v_{\alpha 0}}{c} \zeta_{\alpha}\right)
$$

A1so

$$
A \zeta_{\alpha}=\frac{\gamma_{\alpha}}{\gamma_{\alpha}^{\prime \prime}} A \zeta_{\alpha}^{\prime \prime}
$$

Therefore

$$
\mathrm{d} \xi_{\alpha}^{\prime \prime} \mathrm{d} \eta_{\alpha}^{\prime \prime} d \zeta_{\alpha}^{\prime \prime} / \gamma_{\alpha}^{\prime \prime}=d \xi_{\alpha}{ }_{\alpha} \eta_{\alpha} / \zeta_{\alpha}
$$

Further

$$
N_{\alpha 0} F_{\alpha n}\left(\xi_{\alpha}^{\prime \prime}, \eta_{\alpha}^{\prime \prime}, \zeta_{\alpha}^{\prime \prime}\right)=N_{\alpha} F_{\alpha}\left(\xi_{\alpha}, \eta_{\alpha}, \zeta_{\alpha}\right)
$$

whore $N_{\alpha 0} F_{\alpha n}$ is the enuj.librium dirtributinn function in $S_{\alpha}^{*}$. Alao note that

$$
N_{\alpha}=\gamma_{\alpha 0} N_{\alpha 0}
$$

The expression for $V_{\alpha}$ then becomes

$$
\begin{align*}
& v_{\alpha}=\iiint_{-\infty}^{\infty}\left\{1+\left(\xi^{\prime \prime}+\lambda_{\alpha} \lambda_{x}\right)^{2}+\left(\eta_{\alpha}^{\prime \prime}+\lambda_{\alpha} A_{y^{\prime}}\right)^{2}+\right. \\
& \left.\left[\gamma_{\alpha=}\left(\zeta_{\alpha}^{\prime \prime}+\frac{\gamma_{\alpha 0}}{c} \gamma_{\alpha}^{\prime \prime}\right)+\lambda_{\alpha} A_{z}\right]^{2}\right\}^{\frac{1}{2}} x \\
& \left(1+\frac{v_{\alpha 0} \zeta_{\alpha}^{\prime \prime}}{c r_{\alpha}^{\prime \prime}}\right){ }_{\alpha 0}\left(\xi_{\alpha}^{\prime \prime}, \eta_{\alpha}^{\prime \prime}, \zeta_{\alpha}^{\prime \prime}\right) d \xi_{\alpha}^{\prime \prime} d \eta_{\alpha}^{\prime \prime} d \zeta_{\alpha}^{\prime \prime} \tag{32}
\end{align*}
$$

Now exnarding the coefficient of $F_{\alpha n}$ in the intearanc as a nower series in $\xi_{\alpha}^{\prime \prime} \|_{\alpha}^{\prime \prime} \zeta_{\alpha}^{\prime \prime}$ and then nerforming intecration term by term, assumina $F_{\alpha 0}$ isntronic, ve ohtain

$$
\begin{equation*}
v_{\alpha}=\Lambda_{\alpha}+\frac{{ }_{\alpha}}{2 \Lambda_{\alpha}}\left(1+\alpha_{\alpha 0}\left(\gamma_{\alpha 0}+\frac{5 v_{\alpha n}}{c} L_{\alpha}\right)+\frac{1-u_{\alpha 0}^{2} L_{\alpha}^{2}}{\Lambda_{\alpha}^{2}}\right) \tag{33}
\end{equation*}
$$

where

$$
\begin{align*}
\theta_{\alpha} & =\iiint_{-\infty}^{\infty}\left(\xi_{\alpha}^{\prime \prime 2}, \eta_{\alpha}^{\prime \prime 2}, \zeta_{\alpha}^{\prime 2}\right) F_{\alpha 0} \text { dE } \alpha_{\alpha}^{\prime \prime} \eta_{\alpha}^{\prime \prime} \text { तr, } \alpha_{\alpha}^{\prime \prime} \\
& \equiv \frac{K T_{\alpha}}{m_{\alpha}^{\mathrm{C}^{2}}} \tag{31}
\end{align*}
$$

Note that we have truncated the sorier at the runciratic terms, ignorina bighne order effocte, Nlen $\Lambda_{\alpha}$ is the zoro order term, Which is the result of the cold nlasma. Further, on difforenti?t. the oquation (33) we obtaif

$$
\begin{equation*}
\frac{1}{\lambda_{\alpha}^{2} A} \frac{\partial V^{\alpha}}{\partial A}=\frac{1}{\Lambda_{\alpha}}-\frac{{ }_{\alpha} \alpha}{2 \Lambda_{\alpha}^{3}}\left(1+\gamma_{\alpha 0}\left(\gamma_{\alpha 0}+\frac{5 v_{\alpha n}}{c} L_{\alpha}\right)+\frac{3\left(1-L_{\alpha}^{2} 1^{2}{ }_{\alpha}\right)}{\Delta_{\alpha}^{2}}\right) \tag{35}
\end{equation*}
$$

$$
\begin{align*}
\frac{1}{\lambda_{\alpha}} \frac{\partial V}{\partial A_{z}} & =\frac{L_{\alpha}}{\Lambda_{\alpha}}-\frac{{ }_{\alpha}}{2 \Lambda_{\alpha}}\left(-5 u_{\alpha n}+\frac{\left(\partial \gamma_{\alpha)}^{2}+5 u_{\alpha n} L_{\alpha}-1\right) L_{\alpha}}{\Lambda_{\alpha}^{2}}\right. \\
& \left.+\frac{3\left(1-u_{\alpha \cap}^{2} L_{\alpha}^{2}\right) L_{\alpha}}{\Lambda_{\alpha}^{4}}\right) \tag{35}
\end{align*}
$$

Sincen the analyaie ic corract anl: to the linear terms in $\theta_{\alpha}$, it is normissihle to substituto for $A_{3}$ in the comofficient of ${ }^{A} \alpha$ in the erurtions (35) and (35) the exnres ion aiven by cold nlama results i.e. fruntions (2?) and (30). With these annroximations, the abover eruations become

$$
\begin{align*}
& \frac{1}{\lambda_{\alpha}^{2}} \Lambda \frac{\partial V_{\alpha}}{\partial A}=\frac{J}{\Delta_{\alpha}}-D_{\alpha} \theta_{\alpha}  \tag{37}\\
& \frac{1}{\lambda_{\alpha}}-\frac{\partial V}{\partial \lambda_{z}}=\frac{L_{\alpha}}{\Delta_{\alpha}}-0_{\alpha} \theta_{\alpha}
\end{align*}
$$

where

$$
\begin{align*}
& n_{\alpha}=\frac{1}{2\left(1+\lambda_{\alpha}^{2}\right.} \frac{1}{\left.\Lambda^{2}\right)\left(1+\Omega^{2}\right)^{3 / 2}}\left(5 u_{\alpha \Omega} \cap+\frac{\left(1+\gamma_{\alpha \rho}^{2}\right)\left(1+\Omega^{2}\right)-3 u_{\alpha}^{2} \Omega_{0}^{2}}{\left(1+\Omega^{2}\right)\left(I+\lambda_{0}^{2} A^{2}\right)^{1 / 2}}\right. \\
& \left.+\frac{3}{\left(1+n^{2}\right)\left(1+\lambda: \lambda^{2}\right)^{3 / 2}}\right)  \tag{30}\\
& Q_{\alpha}=\frac{1}{2\left(1+\lambda_{2}^{2} A^{2}\right)^{1 / 6}\left(1+\Omega^{2}\right)^{3 / 2}}\left(-5 n_{\alpha 0}\right. \\
& \left.+\frac{\left[\left(3 \gamma_{\alpha n}^{2}-1\right)\left(1+\Omega^{2}\right)-3 n^{2} \alpha_{0}^{2} \Omega^{2}\right] \Omega}{\left(1+\Omega^{2}\right)\left(1+\lambda_{\alpha}^{2} \Lambda^{2}\right)^{\frac{1}{2}}} \frac{(\Omega}{\left(1+n^{2}\right)\left(1+\lambda_{\alpha}^{2} \Lambda^{2}\right)^{3 / 2}}\right)
\end{align*}
$$

Now substitutind the abo re eruations in (15) and (16) we obtain

$$
\begin{equation*}
\left.\frac{L_{e}}{\Lambda_{e}}=\frac{L_{i}}{\Lambda_{i}}+\partial_{e} \theta_{e}-O_{i} \theta_{i}\right) \tag{A}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{1}{\Lambda_{e}}+\frac{\mu}{\Lambda_{i}}-\left(p_{e}\left(\dot{b}+\mu r_{i} A_{i}\right)=\frac{\omega^{2}}{(1)^{2}} .\right. \tag{12}
\end{equation*}
$$

The next ster $i$ - to eliminate $i_{z}$ sn as to obtain $w$ in torms of the amnliture of the rave $A$ anly. In the circumstance that: the vaves are larae amnlitnde, this is achiovert hy somaring (41), using (?3) are oontinuinc to port only to the lincar termein $\theta_{\alpha}$. Ater enme alonfra mo olitain

$$
\frac{1}{\Lambda_{e}}+\frac{\mu}{\Lambda_{i}}=-\frac{1}{\left(1+\frac{\Lambda^{2}}{2}\right)\left(1+\lambda_{i}^{2} \wedge^{2}\right)}{ }_{i}^{2}\left(O_{i} A_{0} \cdot n_{i} A_{i}\right)
$$

Where we have ascumer $A^{2}>1$ i.f., the rove aro =trone wavor. Using the cole nlasma arnzoseion for $L_{i}$ or ort

$$
\begin{equation*}
\frac{1}{\Lambda_{e}}+\frac{\mu}{\Lambda_{i}}=\left(1+\lambda j \Delta^{2}\right)^{-1} \Omega\left(n_{i} \theta_{i}-n_{c} \theta_{0}\right) \tag{03}
\end{equation*}
$$

so that

$$
\begin{align*}
\frac{\omega^{2}}{\omega_{e}^{2}}= & \left(\frac{n \Omega_{i}}{\left(1+\lambda_{\epsilon}^{2} \wedge^{2}\right)^{2 / 2}} \cdot \mu p_{i}\right) \theta_{i} \\
& -\left(\frac{\Omega O_{0}}{\left(1+\lambda_{\varepsilon}^{2} 1^{2}\right.}+D_{\theta}\right) \theta_{\theta}
\end{align*}
$$

The disnersion mation in $G^{\prime}$ is nbtainat hy using (18)

$$
n^{2}=-1-\Gamma \frac{\omega_{0}^{-2}}{\omega^{2}}\left\{\left(\frac{n_{i}}{\left(1+\lambda_{0}^{2} A^{2}\right)^{\frac{1}{n}}}-\mu D_{i}\right)^{0}{ }_{i}-\left(\frac{0_{e}^{0}}{\left(1+\lambda_{e}^{2} \lambda^{2}\right)^{\frac{1}{3}}}\right.\right.
$$

Where $\Gamma=\left(1-n^{2}\right)^{-1 / 2}$

$$
\left.\left.+\rho_{0}\right\}{ }_{\theta}\right\}
$$

From the nemrossion of $Q_{\alpha}$ and $\sigma_{\alpha^{\prime}}$. it $i=$ evidrat that $n_{i}$ and $P_{i}$ can he larco comnared to $n$ and $n$ romoctivoly. We may thex?fore conclude that the innic contributions can he airnificant unloss the inn-tomnoratur is nealirifly snall.


In this section we trost the amnlituce nf the wave as a smill marameter, and then use the norturbation techniaun th determine the disncesion relation inenemontint first arer non-linear correction, tut the temnerature in this canc is unrestricted. To to explicit, we ehall expane ${ }_{x}$ in nowers of $A$ and $A$ and then truncate the serins at trime nf nrider $A^{3}$. With this $V_{\alpha}$, we calculate its तifforontials $\frac{\partial V \alpha}{\partial A}$ and $\frac{\partial V_{\alpha}}{\partial \pi_{z}}$ an' then substituter them in ocuations (15) and (16). That mill virirt the desired dismorsinn rolation.

For ennvenience wo uen the cylinerical noler cocreinnes $\underline{u}_{\alpha}=\left(\rho_{\alpha}\right.$ ens $\phi_{\alpha}, \rho_{\alpha}$ rin $\left.\phi_{\alpha}, \zeta_{\alpha}\right)$ and adのnt the netation

$$
\left\langle P\left(\rho_{\alpha}, \phi_{\alpha}, \zeta_{\alpha}\right)\right\rangle=\int D\left(0_{\alpha}, \phi_{\alpha}, \zeta_{\alpha}\right) F_{\alpha}\left(\underline{U}_{\alpha}\right){\underset{\beta}{ }}^{3} \underline{u}_{\alpha}
$$

$$
\begin{align*}
\frac{1}{D_{\alpha}}= & \frac{\lambda}{\gamma_{\alpha}}\left(1-\frac{1}{\gamma_{\alpha}^{2}} \lambda_{\alpha} \Lambda_{\alpha} \cos \phi_{\alpha}-\frac{1}{2 \gamma_{\alpha}^{2}}\left(\lambda_{\alpha}^{2} A^{2}+2 \lambda_{\alpha} \lambda_{z} \zeta_{\alpha}\right)\right. \\
& +\frac{3}{2 \gamma_{\alpha}^{4}} \lambda_{\alpha}^{2} A^{2} \rho_{\alpha}^{1} \cos ^{2} \phi_{\alpha}+\frac{3}{2 \gamma_{\alpha}^{4} \lambda_{\alpha}} A_{\alpha} \operatorname{cns} \phi_{\alpha}\left(2 \lambda_{\alpha} A_{z} \zeta_{\alpha}\right. \\
& \left.\left.+\lambda_{\alpha}^{2} \lambda^{2}\right)-\frac{5}{2 \gamma_{\alpha}^{6}} \lambda_{n}^{3} \Lambda^{3} \rho_{\alpha}^{3} \cos ^{3} \phi_{\alpha}\right) \tag{19}
\end{align*}
$$

where we have assumed that $\lambda_{z}$ is of the order of $A^{2}$. This assumption is indeed true for cold nlusma, as may he seen from the equation (28) and is verified n nasterireri for the hat plasma (see oruation51).

$$
\begin{align*}
& \text { Hone, to the ort ar } A^{2} \\
& \frac{1}{\lambda_{\alpha}} \frac{\partial_{\alpha} \eta_{\alpha}}{\partial \Lambda_{z}}=<\frac{\lambda_{n} \Lambda_{z}}{\gamma_{\alpha}}+\frac{\gamma_{\alpha}}{\gamma_{\alpha}}\left(1-\frac{\lambda_{\gamma}^{2} \Lambda^{2}+2 \lambda_{\alpha} z^{\zeta} \alpha}{2 \gamma_{\alpha}^{2}}\right. \\
& \left.+\frac{3 \lambda_{\alpha}^{2} \Lambda^{2} \rho_{1}}{Y_{\alpha}^{4}}\right)> \tag{53}
\end{align*}
$$

Here wo necrose that

$$
\left\langle\frac{\zeta_{i}}{\gamma_{i}}\right\rangle=\left\langle\frac{\zeta_{e}}{\gamma_{e}}\right\rangle
$$

because from the emulation (J), it is clear that

$$
J_{C}+\left.T_{i}\right|_{\text {at } A=?}=?
$$

1.e.

$$
\int \frac{u_{0}}{\gamma_{r}} F_{C}\left(\underline{u}_{C}\right) d^{3} \underline{u}_{c}=\int \frac{\underline{u}_{i}}{\gamma_{i}} F_{i}\left(\underline{u}_{i}\right) d^{3} \underline{u}_{i}
$$

Therefore the relation (15) determince $z_{z}$ as

$$
\lambda_{e} A_{z}=\frac{\left\langle\frac{\zeta_{G}}{2 \gamma \Omega}\left(1-\frac{3 \rho_{0}^{2}}{2 \gamma^{3}}\right)-\frac{\mu^{2} \zeta_{i}}{2 \gamma 1}\left(\lambda-\frac{30_{i}^{2}}{2 \gamma^{2}}\right)\right\rangle}{\left\langle\frac{1+\rho_{0}^{2}}{\gamma_{0}^{3}}+\mu \frac{1+D_{1}^{2}}{\gamma\}}\right\rangle} \lambda^{2} e^{A^{2}}
$$

Also to the oricer $A^{2}$, it is found that.

$$
\begin{align*}
\frac{1}{\lambda_{\alpha}^{2} A} \frac{\partial V}{\partial A}= & \left\langle\frac{1}{\gamma_{\alpha}}-\frac{\rho_{\alpha}^{2}}{2 \gamma_{\alpha}^{3}}-\frac{\zeta_{\alpha}}{\gamma_{\alpha}^{3}}\left(1-\frac{3 \rho_{\alpha}^{2}}{2 \gamma_{\alpha}^{2}}\right) \lambda_{\alpha} \lambda_{z}\right. \\
& \left.-\frac{1}{2 \gamma_{\alpha}^{3}}\left(1-\frac{3 \rho_{\alpha}^{2}}{\gamma_{\alpha}^{2}}+\frac{15}{8} \frac{\rho_{N}^{4}}{\gamma_{\sim}^{4}}\right) \lambda_{\alpha}^{2} A^{2}\right\rangle \tag{5}
\end{align*}
$$

Now using (51) and (52) in ermation (1r) we att

$$
\begin{equation*}
\frac{\omega^{2}}{\omega_{2}^{2}}=x_{0}-\frac{3}{2} X_{1} \lambda_{6}^{2} A^{2} \tag{5.3}
\end{equation*}
$$

phere

$$
x_{0}=\left\langle\left(\frac{1}{\gamma_{e}}-\frac{\rho_{0}^{2}}{2 \gamma_{e}^{3}}\right)+\mu\left(\frac{1}{\gamma_{i}}-\frac{\sigma_{i}^{2}}{2 \gamma_{i}^{3}}\right)\right\rangle
$$

whore

$$
\begin{aligned}
\mathrm{x}_{1}= & \frac{\left\langle\frac{\zeta_{C}}{\gamma_{C}^{3}}\left(1-\frac{3 \rho_{c}^{2}}{2 \gamma_{C}^{2}}\right)-\frac{\mu^{2} \zeta_{i}}{\gamma_{1}^{3}}\left(1-\frac{3 \rho_{i}^{2}}{2 \gamma_{i}^{2}}\right)^{2}\right.}{\left.1+\frac{\rho_{c}^{2}}{\gamma_{e}^{3}}+1-\frac{\rho_{i}^{2}}{\gamma_{1}^{3}}\right\rangle}+ \\
& \left\langle\frac{1}{\gamma_{e}^{3}}\left(1-\frac{3 \rho_{c}^{2}}{\gamma_{C}^{2}}+\frac{15}{9} \frac{\rho_{c}^{4}}{\gamma_{e}^{4}}\right)+\frac{\mu^{3}}{\gamma_{i}^{3}}\left(1-\frac{3 \rho_{i}^{2}}{\gamma_{i}^{2}}+\frac{15}{8} \frac{\rho_{i}^{4}}{\gamma_{i}^{3}}\right)\right\rangle
\end{aligned}
$$

Note that the effecte of the ionic motion stan out in the co-effici nt of $\mu$ anc? takina $\mu=0$ rives the ole results of the one-comoonent nlasma.

$$
\text { In the frame } s^{\prime} \text { the }{ }^{2} \text { isnersion relation takes tho }
$$

form:

$$
\begin{equation*}
n^{2}=1-\Gamma\left(x_{n}-\frac{1}{2} y_{1} \lambda_{2}^{2} A^{2}\right)\left(\frac{\omega^{\prime}}{\omega^{\prime}}\right)^{2} \tag{56}
\end{equation*}
$$

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Recently Ichikawa et al ${ }^{(1)}$ studied the old 'recurrence problem' of Fermi, Pesta and Ulam of one-dimensional annarnonic lattice and gave an explanation emphasizing the discrete ciaracter of the system in terms of phonons, in contrast to the Jabusisy's Continuum Model. (2) In their work, the Korteveg-de Vries (IdV) ecuatiol is derived on the basis of a coherent state representation for the interacting phonons, and it is explicitly shown that a soliton solution can be given a quantum mechanical interpretation as a coherent state of excited phonons in the system. In the present note, we extend this worl to a generalized one-dimensionz? lattice with an arbitrary degree of anharmonicity $n$ and obtain a generaiized "idy equation that describes the system. The cuantur conecnt of the onesoliton state is męintained as before. An expression for the offective mass of the soliton is also given in terms of the degree $n$ and the coupling $g_{n}$ of non-linearity. In this work, se shall mostly follow the notation of reference (1).

We consider a one-dimensional generalized anharmonic
lattice with N-particles equally spaced over a length $\mathrm{L}=\mathrm{H}$ ( ciescribed by the Ramiltonian

$$
\begin{gather*}
H=\sum_{r=1}^{N} \frac{1}{2}\left[m \dot{y}_{r}^{2}+k\left(y_{r+1}-y_{r}\right)^{2}+\frac{1}{n} k g_{n}\left(y_{r+1}-y_{r}\right)^{n}\right]_{0} \\
(n=3,4,5, \cdots) \tag{i}
\end{gather*}
$$

Aore $y_{r}, \dot{y}_{r}$ are the displacement and velocity of the rti? particle itin mass $m, K$ is the linear spring constant and $g_{n}>0$ measures he strength of the non-linearity. Introducing the nomal mode xpansions

$$
\begin{align*}
& y_{r}=\frac{1}{\sqrt{N}} \sum_{k} \sqrt{\frac{k}{2 m \omega(k)}}\left(a_{-k}^{*}+a_{k}\right) e^{i k x_{r}}  \tag{2}\\
& \dot{y}_{r}=\frac{i}{\sqrt{N}} \sum_{k} \sqrt{\frac{k \omega(k)}{2 r}}\left(a_{-k}^{*}-a_{k}\right) e^{i k x_{r}}
\end{align*}
$$ where $x_{r}=r l$ gives the position of the ruth particle. Now quantiziag the system in the usual way, we obtain

$$
\begin{align*}
H= & H_{0}+H^{\prime} \\
H_{0}= & \sum_{k} \hbar \omega(k)\left(a_{k^{\prime}}^{*}+1 / 2\right) \\
H^{\prime}= & \sum_{k_{1}} \Delta\left(k_{2}, \cdots, k_{n}+k_{2}+\cdots+k_{n}\right) \varphi\left(k_{1}, k_{2}, \cdots, k_{n}\right) \times \\
& x\left(a_{-k_{1}}^{*}+a_{k_{1}}\right)\left(a_{-1_{2}}^{*}+a_{k_{2}}\right) \cdots\left(a_{-k_{n}}^{*}+a_{k_{n}}\right) \tag{3}
\end{align*}
$$

where

$$
\begin{align*}
\Delta(k) & =\frac{1}{N} \sum_{r=1}^{N} e^{i r l k}  \tag{4}\\
\omega^{2}(k) & =4 \frac{K}{m} \sin ^{2}\left(\frac{l k}{2}\right)  \tag{5}\\
\varphi\left(k_{1}, k_{1}, \cdots, k_{n}\right) & =\frac{1}{2}\left(\frac{1}{n_{1}} K g_{n}\right)\left(\frac{k}{2 m}\right)^{r / 2} \frac{(2 i)^{n}}{(\sqrt{N})^{n-2}} \exp \left\{-\frac{i l}{2}\left(k_{1}+\cdots+k_{n}\right)\right\} x \\
x & \left\{\omega\left(k_{1}\right) \omega\left(k_{2}\right) \ldots \omega\left(k_{1}\right)\right\}^{-1 / 2} \sin \frac{l k_{1}}{2} \sin \frac{l k_{2}}{2} \ldots \sin \frac{l k_{n}}{2}
\end{align*}
$$

And

$$
\left[a_{k}, a_{k^{\prime}}^{*}\right]=\Delta\left(k-k^{\prime}\right) ;\left[a_{k}, i_{k}\right]=0=\left[a_{k}^{*}, a_{k^{\prime}}^{*}\right]
$$

Now we introduce, following Glauber (3), the coherent state of phonons $\left|\alpha_{k}\right\rangle$ defined as:

$$
\begin{aligned}
a_{k}\left|\alpha_{k}\right\rangle & =\alpha_{k}\left|\alpha_{k}\right\rangle \\
\left|\alpha_{k}\right\rangle & =\exp \left(-1 / 2\left|\alpha_{k}\right|^{2}\right) \sum_{n_{k}=0}^{\infty} \frac{\left(\alpha_{k}\right)^{n}}{\sqrt{n_{k}!}}\left|n_{k}\right\rangle
\end{aligned}
$$

with average occupation number given by a Poisson distribution ifitw mean value $\left\langle n_{k}\right\rangle=\left|\alpha_{k}\right|^{2}$. Then the expectation value of the dis?acement is given by
where

$$
\begin{aligned}
\left\langle\alpha_{k}\right| y_{r}\left|\alpha_{k}\right\rangle & =\frac{1}{\sqrt{N}} \sum_{k} y(k) e^{i k x_{r}} \\
y(k) & =\sqrt{\frac{k}{2 m \omega(k)}}\left(\alpha_{-k}^{*}+\alpha_{k}\right)
\end{aligned}
$$

Using the temporal evolution of the expectation values of the Heiwonjes creation and destruction operators with respect to a coherent state, we obtain the equation of motion for the kith mode displacement $y(=)$ :

$$
\ddot{y}(k)=-\omega^{2}(k) y(k)-2 n \sqrt{\frac{\omega(k)}{2 m-k}} \cdot\left(\frac{2 m}{\hbar}\right)^{\frac{n-1}{2}} \sum_{k_{1}, \ldots, k_{n}} \varphi\left(k_{1}, \ldots, k_{n}\right) x
$$

$$
x \Delta\left(k_{1}+\cdots+k_{n}\right) \Delta\left(k+k_{1}\right) \sqrt{\omega\left(k_{1}\right) \ldots \omega\left(k_{n}\right)} y\left(k_{2}\right) \ldots y\left(k_{n}\right)
$$

If we neglect the contributions from the large wave-number phono:2s, we may approximate $\phi$ and $\omega$ (egns. (5) and (6)) as

$$
\begin{aligned}
\phi\left(k_{1}, \ldots, k_{n}\right)= & \frac{1}{2}\left(\frac{1}{n} k g_{n}\right)\left(\frac{k}{2 m}\right)^{n / 2} \cdot \frac{(2 i)^{n}}{(\sqrt{N})^{n-2}} \cdot \frac{l k_{1}}{2} \cdot \frac{l k_{2}}{2} \cdots \cdot \frac{l k_{n}}{2} x \\
& x\left\{\omega\left(k_{1}\right) \cdot \omega\left(k_{2}\right) \cdots \omega\left(k_{n}\right)\right\}^{-1 / 2} \\
\omega(k) & \approx \sqrt{\frac{k}{m}} l|k|\left(1-\frac{1}{24} l^{2} k^{2}\right) \\
& =s|k|\left(1-\frac{1}{24} l^{2} k^{2}\right), \text { where } s=\sqrt{\frac{k}{m}} \ell
\end{aligned}
$$

that the equation of motion (7) becomes

$$
\begin{align*}
& \ddot{y}(k)+s^{2} k^{2}\left(1-\frac{1}{12} e^{2} k^{2}\right) y(k)=\frac{1}{8} k s^{2}\left(\frac{l}{2 \sqrt{N}}\right)^{n-2}(2 i)^{n} \cdot k x \\
& x \sum_{k} \Delta\left(-k+k_{2}+\cdots+k_{n}\right) k_{2} k_{3} \ldots k_{n} y\left(k_{2}\right) \cdots y\left(k_{n}\right) \tag{3}
\end{align*}
$$

w defining a new variable $u(k, t)$ and its Fourier transform

$$
\begin{aligned}
& u(k, t)=i k y(k, t) \\
& u(x, t)=\frac{1}{\sqrt{N}} \sum_{k} u(k, t) e^{i k x}
\end{aligned}
$$

may Fourier transiorm Eqn. (8) into a non-linear cificiential nation which governs the dynamics of our generalized aminarmonic trice:

$$
\begin{array}{r}
\frac{\partial^{2}}{\partial t^{2}} u(x, t)-s^{2} \frac{\partial^{2}}{\partial x^{2}} u(x, t)-\frac{1}{12} s^{2} e^{2} \frac{\partial^{4}}{\partial x^{4}} u(x, t)-g_{n} \frac{s^{2} e^{n-2}}{2} x \\
x \frac{\partial^{2}}{\partial x^{2}}(u(x, t))=0 \tag{9}
\end{array}
$$

is equation is a ficneralization of the Boussinesc ervation and In be converted to the lid type by using the reductive perturbation
method (4) with the following espansion and soace-tine rescaling

$$
\begin{aligned}
& u=\epsilon u^{(1)}+\epsilon^{2} u^{(2)}+ \\
& \xi=\epsilon^{\left(\frac{n-2}{2}\right)}(x-t) \\
& \tau=\epsilon^{3\left(\frac{n-2}{2}\right)} t
\end{aligned}
$$

into the form

$$
\frac{\partial}{\partial \xi}\left[\frac{\partial}{\partial \tau} u^{(1)}+\frac{1}{24} s^{2} l^{2} \frac{\partial^{3}}{\partial \xi^{3}} u^{\prime}+g_{n} \frac{s^{2} e^{n-2}}{4} \frac{\partial}{\partial \xi}\left(u^{(1)}\right)^{n-1}\right]=0
$$

(10)

- the generalized iidV equation. Hotice that ow cacosing $n=3,4$ onerecovers the standard KdV erration and its lociried form for cab:e and quartic non-linearities, re:pectively. Zeturnias to the origin.... variables, the above ecuation ( 00 ), becomes

$$
\begin{equation*}
\frac{\partial}{\partial t} u(x, t)+\frac{\partial}{\partial x} u(x, t)+\frac{1}{24} s^{2} e^{2} \frac{\partial^{3}}{\partial x^{3}} u(x, t)+g_{n} \frac{s^{2} e^{r-2}}{2} \frac{\partial}{\partial x}[u(x, t)]^{r-1}=0 \tag{11}
\end{equation*}
$$

The goneralized "idy ecuation acm ts one soliton zoletion winch is given by ${ }^{(3)}$

$$
\begin{equation*}
u(x, t)=\eta\left[\operatorname{sech} a_{1}\left(x-b_{n} s t\right)\right]^{\frac{2}{n-2}} \tag{12}
\end{equation*}
$$

where

$$
\begin{aligned}
& a_{n}^{2}=(n-2)^{2}\left(\frac{3}{n} g_{n} \eta^{n-2} e^{n-4}\right) \\
& b_{n}=\left(1+\frac{1}{2 n} g_{n} \eta^{n-2} e^{n-2}\right)
\end{aligned}
$$

It is easy to see that $n=3$ renroduces the resizts of ichikawa et ${ }_{a}{ }^{(1)}$.

Finally, it is strrig teorvard to verify jut
 statie of excited rio ons wit: amilitece $\alpha_{l}$ as

$$
\begin{array}{r}
a_{k}=\frac{-i}{\sqrt{N}} \int^{\frac{m \omega(k)}{2 k}}\left(1+b_{n} \frac{s k}{\omega i k}\right) 4^{1 / n-2} \cdot\left(\frac{\eta}{2 a_{n} 2 k}\right) \times \\
\quad \times e^{-i k b_{n} s t} B\left(\frac{1}{n-2}+\frac{i k}{2 a_{r}} ; \frac{1}{n-2}-\frac{i k}{2 a_{n}}\right)
\end{array}
$$


Gere the beta finaction $3(\mu, \nu)$ enteas tirroun an fovider transforn of ite one-soition solution:

$$
\begin{aligned}
u(k, t) & =\frac{\sqrt{N}}{L} \int_{-\infty}^{\infty} u(x, t) e^{-i k x} c x \\
& =\frac{4 / n-2}{\sqrt{N} 2 a_{n} l} e^{-i k b_{r} s t} E\left(\frac{1}{n+2}+\frac{i k}{2 a_{n}}, \frac{1}{n-2}-\frac{i k}{2 a_{n}}\right)
\end{aligned}
$$


 ation

$$
\begin{align*}
\underline{P} & =\sum_{k} h \underline{k}\left\langle n_{k}\right\rangle \\
& =m_{n}^{*}\left(1+\frac{\alpha_{n}}{\partial n} \eta^{n-2} l^{n-2}\right) \underline{s} \tag{14}
\end{align*}
$$

wherc $n_{n}$ ．cefinec as the effective ass $0=32$ aniton is siven bjo

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